

Gravitational Ep and infinity

In past $E_p = mgh$. No more. Because g changes as distance from a mass changes the definition of E_p must also change.

$E_p = mgh$ works ONLY when the entire situation occurs close to a planet's surface.

The Law of Conservation of Energy still applies:

$E_{p0} + E_{k0} = E_{pf} + E_{kf} + E_h$ but

E_p is NOT mgh any more.

~~BAD~~

$E_p \text{ at } \infty = 0$

E_p closer than ∞ must be < 0

We need a new place to define E_p , that place is infinity \leftarrow infinitely far from all masses.

E_p will be - or zero

The gravitational potential energy of some object (m_1) is zero Joules when infinitely far from some planet (m_2)

$E_p = -\frac{Gm_1m_2}{r}$

All E_p measured in relation to infinity will be less than zero (negative).

$E_p = -\frac{Gm_1m_2}{r}$ r is the distance to the centre of the mass (planet).

distance from centre

* Calculate your E_p , relative to infinity at the surface of the moon if your mass is

100 kg.

$E_p = -\frac{Gm_1m_2}{r}$

$= -\frac{6.67 \times 10^{-11} (100) 7.35 \times 10^{22}}{1.74 \times 10^6} = -2.81 \times 10^8 \text{ J}$

You are now lifted to a radius of $8.0 \times 10^9 \text{ m}$ from the moon's centre. What Was the work done?

$W = \Delta E_p$
 $W = E_{\text{final}} - E_{\text{initial}}$

W - work

$$W = E_{pf} - E_{po}$$

$$= -\frac{Gm_1m_2}{r_f} - \frac{Gm_1m_2}{r_o}$$

$$W = -\frac{6.67 \times 10^{-11} \cdot 100 \cdot 7.35 \times 10^{22}}{8 \times 10^9} - \frac{6.67 \times 10^{-11} \cdot 100 \cdot 7.35 \times 10^{22}}{1.74 \times 10^6}$$

$$W = \frac{-61280}{2.82 \times 10^8} + \frac{2.82 \times 10^8}{2.82 \times 10^8} = 2.82 \times 10^8 \text{ J}$$

An asteroid of mass 1×10^4 kg can be considered infinitely far from Earth. It is moving through space at 1 km / s & heads to Earth, with what speed will it impact us?

~~$$E_{po} + E_{ko} = E_{pf} + E_{kf}$$~~
~~$$\frac{1}{2} m v_o^2 = -\frac{Gm_1m_2}{r_o} + \frac{1}{2} m v_f^2$$~~

$$\frac{1}{2} 1000^2 = -\frac{6.67 \times 10^{-11} \cdot 5.98 \times 10^{24}}{6.38 \times 10^6} + \frac{1}{2} v_f^2$$

$$500000 = -6.25 \times 10^7 + \frac{1}{2} v_f^2$$

$$\sqrt{\frac{6.30 \times 10^7}{\frac{1}{2}}} = \frac{1}{2} v_f = 1.1 \times 10^4 \frac{m}{s}$$

Escape Velocity: this is the minimum velocity to escape a planet's gravitational pull. **This means that you make it to infinity, but are at rest when you get there.**

surface ∞

$$E_{p0} + E_{k0} = E_{pf} + E_{kf} + \cancel{E_{th}}$$

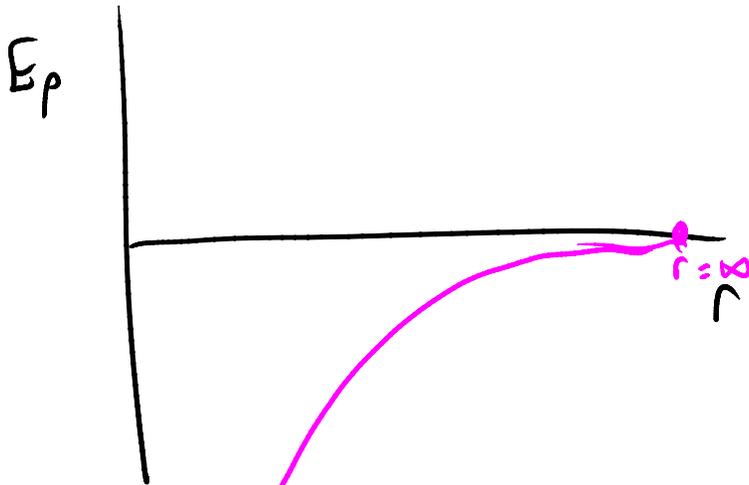
$$\frac{-Gm_0m_e}{r_e} + \frac{1}{2}m_0v_0^2 = 0 + 0$$

$$\frac{1}{2}m_0v_0^2 = \frac{Gm_0m_e}{r_e}$$

$$v_0 = \sqrt{\frac{2 \left(\frac{6.67 \times 10^{-11} \cdot 5.98 \times 10^{24}}{6.38 \times 10^6} \right)}{2}} = 1.12 \times 10^4 \frac{m}{s}$$

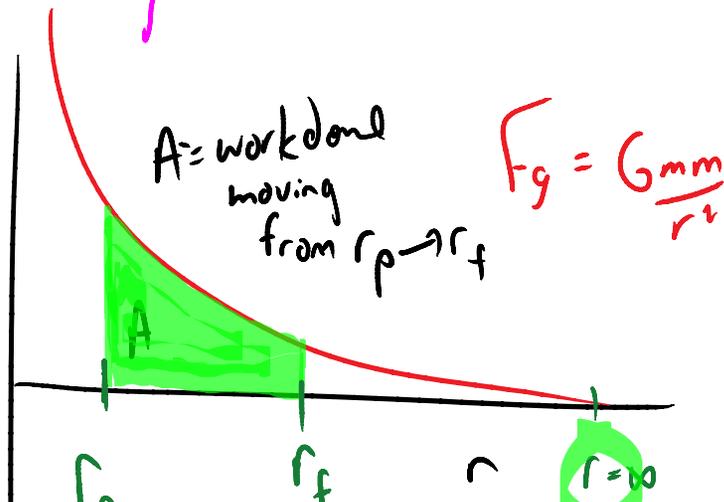
Graphs of E_p vs. r

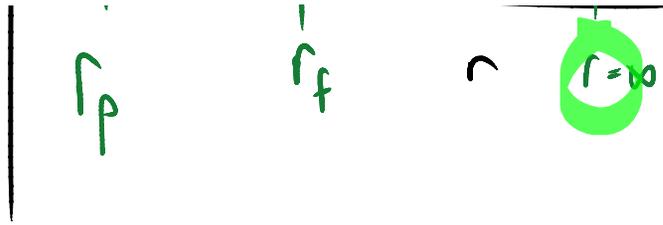
$$E_p = -\frac{Gmm}{r}$$



Graphs of F_g vs. r

$$F_g$$





Escape: rockets like to escape the planet they are on.

If you escape you must be free of a gravitational field g , and you can stop.
This means $E_{pf} = 0$ and $E_{kf} = 0$

So:

By the Law of Conservation of Energy:

$$E_{po} + E_{ko} = E_{pf} + E_{kf} + Q$$